Numerical simulation of the improved dynamics of sedimentary river beds with a stochastic Exner equation

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joint work with

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Outline

Context & Motivations

Saint-Venant-Exner equations

Stochastic case

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Context Organisation

Thesis co-supervised by:

- Université d'Orléans: Stéphane Cordier,
- Université Paris 13: Emmanuel Audusse,
- LHSV, LNHE, EDF R&D: Sébastien Boyaval, Nicole Goutal, Magali Jodeau,

and co-financed by:

- AMIES,
- EDF R&D.

Partnership: Team ANGE

- CEREMA,
- Inria Rocquencourt,
- Université Pierre et Marie Curie: LJLL.

Sediments transport is responsible of modification of river beds. 2 processes of sediments transport:

- by suspension: particles can be found on the whole vertical water depth and rarely be in contact with the bed,
- by bedload: particles are moving near the bed by saltation and rolling.

In the following, we only focuse on the **bedload transport**. In the literature, most of industrial codes use the **Saint-Venant-Exner model**.

Motivations Stochastic viewpoint

Work initiated during the CEMRACS 2013.

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At the grain scale to the laboratory one, physical experiments reveal **fluctuations of the solid flux** (Recking et al.).



This problem has been investigated by physical theory and **various stochastic Exner equations** have been proposed (Jerolmack & Mohrig 2005, Ancey 2010 & 2014, Furbish et al. 2012).

We propose a numerical study of a possible stochastic Exner equation.

Saint-Venant-Exner equations

Coupled model:

$$\int \partial_t H + \partial_x \left(\mathbf{Q} \right) = \mathbf{0}, \tag{1a}$$

$$\partial_t \mathbf{Q} + \partial_x \left(\frac{\mathbf{Q}^2}{H} + \frac{gH^2}{2} \right) = -gH\partial_x B - \frac{\tau}{\rho}.$$
 (1b)

$$\bigcup_{t} \partial_t B + \partial_x Q_s = 0, \tag{1c}$$

between:

 the Saint-Venant equations (aka shallow-water equations): (1a)-(1b)



H(t, x): water height, Q(t, x) = HU: discharge, B(t, x): bottom topography, with $x \in \Omega \subseteq \mathbb{R}$, $t \ge 0$.

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Saint-Venant-Exner equations

 τ is defined by the Manning formula,

$$\tau = \rho g H \frac{Q|Q|}{H^2 \kappa_s^2 R_h^{4/3}},\tag{2}$$

where, in the particular case of a rectangular channel with width I, the hydraulic radius R_h reads

$$R_h = \frac{IH}{I+2H}.$$

Saint-Venant-Exner equations

the Exner equation (1c)

where $Q_s(t,x)$ is the solid transport flux defined by

$$Q_{s} = \sqrt{\frac{g(\rho_{s} - \rho)d^{3}}{\rho}} Q_{s}^{\star}(\tau^{\star}; \tau_{c}^{\star}) \frac{\tau^{\star}}{|\tau^{\star}|}$$
(3)

and the Meyer-Peter-Müller formula,

$$Q_{s}^{\star} = A \left(|\tau^{\star}| - \tau_{c}^{\star} \right)^{3/2}$$
(4)

 $\text{with} \left\{ \begin{array}{ll} A & \text{the characteristic length of a grain jump,} \\ \rho_s, \rho & \text{resp. the mass densities of the solid and fluid phases,} \\ g & \text{the gravitational acceleration,} \\ \tau^* & \text{the shear stress (aka Shields parameter),} \\ \tau_c^* & \text{the critical value for the initiation of motion,} \\ d & \text{the grain diameter.} \end{array} \right.$

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$$V = \sqrt{rac{g(
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ho)d}{
ho}} a(\sqrt{ au^{\star}} - \sqrt{ au_c^{\star}})$$

where V has an exponential distribution and under the local equilibrium $\tilde{n} \propto (\tau^* - \tau_c^*)$, we retrieve the **Meyer-Peter-Müller** formula

$$Q_s^{\star} = A\left(|\tau^{\star}| - \tau_c^{\star}\right)^{\frac{3}{2}}$$

with A a stochastic coefficient following an **exponential** distribution; A(t,x) independent in time and space.

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Model defined on the torus: **Periodic boundary conditions**.

Space discretization, $2 \times N_x$ cells denoted $C_{i, i+1/2}$, $i = 0, \dots, N_x - 1$, $\underbrace{C_i}_{C_{i+1/2}}, \underbrace{C_{i, i+1/2}}_{C_{i, i+1/2}}, \underbrace{C_{i, i+1/2}}_{X_i, x_{i+1/2}}, \underbrace{C_{i, i+1/2}}_{X_{i+1/2}}, \underbrace{C_{i, i+1/2}}_{$

1. Finite Volume method for the fluid quantities

$$W_{i}^{n+1} = W_{i}^{n} - \frac{\Delta t^{n}}{\Delta x} \left(F_{i+1/2}^{n} - F_{i-1/2}^{n} \right) + \frac{\Delta t^{n}}{\Delta x} S(W_{i}^{n}, B_{i-1/2}^{n}, B_{i+1/2}^{n}),$$
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$$F_{i+1/2} = F(W_i, W_{i+1}) = \frac{F(W_i) + F(W_{i+1})}{2} - c \frac{W_{i+1} - W_i}{2},$$
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with $F(W) = (Q, \frac{Q^2}{H} + \frac{gH^2}{2})^T,$
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with $F(W) = (Q, \frac{Q^2}{H} + \frac{gH^2}{2})^T,$
 $c = \max(|U_i| + \sqrt{gH_i}, |U_{i+1}| + \sqrt{gH_{i+1}})$ and
 $S(W_i^n, W_{i+1}^n)$ is a discrete source term written as
 $S(W_i^n, B_{i-1/2}^n, B_{i+1/2}^n) = \begin{pmatrix} 0 \\ gH_i^n(B_{i+1/2}^n - B_{i-1/2}^n) \end{pmatrix}.$ (7)

2. Bottom topography appoximated by a **centered finite difference formula**,

$$B_{i+1/2}^{n+1} = B_{i+1/2}^{n} - \frac{\Delta t^{n}}{\Delta x} \left(Q_{s}(H_{i+1}^{n}, U_{i+1}^{n}) - Q_{s}(H_{i}^{n}, U_{i}^{n}) \right),$$
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Discrete stochastic solid flux given by the stochastic **Meyer-Peter-Müller formula**

$$(Q_{s})_{i}^{n} = \sqrt{\frac{g(\rho_{s} - \rho)d^{3}}{\rho}} A_{i}^{n} (|(\tau^{*})_{i}^{n}| - \tau_{c}^{*})^{\frac{3}{2}} sg((\tau^{*})_{i}^{n})$$

where A_i^n are independent identically distributed random variables with exponential distribution.

Stochastic Saint-Venant–Exner model Monte-Carlo simulations

Description of the deterministic case

- Stationary uniform flow in torrential regime,
- Sloped bottom topography.

Number of realizations M = 1000.

 T_{fin} sufficiently large such that the empirical variance of all the quantities of interest seems close to long-time stationary values.

Monte-Carlo simulations for the SVE model Numerical results



Figure: Empirical mean of the topography deviation for two meshes (left) and empirical variance of the topography deviation for different meshes (right)

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Monte-Carlo simulations for the SVE model Numerical results



Figure: Empirical mean of the velocity for two meshes (left) and empirical variance of the velocity for different meshes (right)

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Figure: PDFs for the bottom topography deviation (left) the velocity (right) for the finest mesh

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Expression of the perturbed topography

$$\begin{cases} B_{i+1/2}^{n} = B_{i+1/2}^{0} + \tilde{B}_{i+1/2}, \\ \tilde{B}_{i+1/2} = \alpha \sqrt{\Delta x} \sum_{k=1}^{N/2} \frac{1}{k} \left(a_{k} \cos\left(2k\pi \frac{i+1/2}{N}\right) + b_{k} \sin\left(2k\pi \frac{i+1/2}{N}\right) \right), \end{cases}$$

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where $B_{i+1/2}^{0}$ corresponds to the non-perturbed initial bottom topography,

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where $B_{i+1/2}^0$ corresponds to the non-perturbed initial bottom topography, a_k and b_k are values obtained with a normal law $\mathcal{N}(0,1)$, and α is an imposed amplitude.

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Figure: PDF for the velocity (left) and empirical mean of the velocity as a function of time (right) with $N_x = 150$

Stochastic Saint-Venant system Calibration with the Strickler coefficient K_s

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Idea

The new coefficient is computed at each time step as a deterministic function of moments of H and Q. The expression is established by imposing that no energy is dissipated by the model,

$$\mathbb{E}\left[\sum_{i=0}^{N_{x}-1} \frac{\boldsymbol{Q}_{i}^{n+1}}{\boldsymbol{Q}_{i}^{n+1}}\right] = \mathbb{E}\left[\sum_{i=0}^{N_{x}-1} \frac{\boldsymbol{Q}_{i}^{n}}{\boldsymbol{Q}_{i}^{n}}\right]$$

Stochastic Saint-Venant system Calibration with the Strickler coefficient K_s

New Strickler coefficient

$$(K_{s})^{n} = \left(-\frac{\mathbb{E}\left[\sum_{i=0}^{N_{x}-1} \frac{Q_{i}^{\omega,n} |Q_{i}^{\omega,n}|}{H_{i}^{\omega,n} (R_{h,i}^{\omega,n})^{4/3}}\right]}{(-\partial_{x}B^{0}) \times \mathbb{E}\left[\sum_{i=0}^{N_{x}-1} H_{i}^{\omega,n}\right] + \frac{1}{\Delta x} \mathbb{E}\left[\sum_{i=0}^{N_{x}-1} H_{i}^{\omega,n} (\tilde{B}_{i+1/2}^{\omega} - \tilde{B}_{i-1/2}^{\omega})\right]}\right)^{1/2}}$$
(9)



Figure: PDF for the velocity (left) and empirical mean of the velocity as a function of time (right) with $N_x = 150$



Figure: Strickler coefficient K_s as a function of time for different meshes

Perspectives

- Modelling the noise introduced in Saint-Venant-Exner so as to compensate for the dissipation introduced: scaling of the injected noise with respect to time and space?
- Ability to maintain the deterministic equilibrium in the mean when the bottom topography is perturbed in time in a stochastic Saint-Venant-Exner.
- Robustness of the new model concerning the convergence to a continuous time-space model.
- Comparison of the physical information supported by the stochastic variable A (characteristic size of ripples...) with other ones (τ*).

Thank you for your attention!